

Looking for seasonal changes in Imhotep with MIRO

Anthony Lethuillier¹, Paul von Allmen¹, Mark Hofstadter¹, and the MIRO team.

¹Jet Propulsion Laboratory/Calif. Inst. Tech., Pasadena, CA, United States.

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Jet Propulsion Laboratory
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Context

Context

1. Context
2. Model of the subsurface
3. Results

The ToO#7:

- First observation: October 27th 2014 as a single swath.

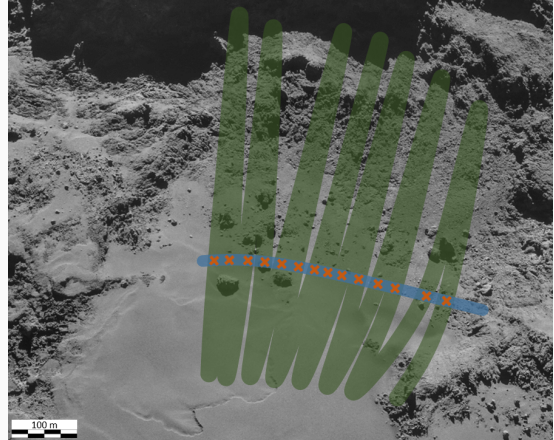


Figure 1: NAVCAM image of the Imhotep region with indicated the 2014 and 2016 swaths. Copyright: ESA/Rosetta/NAVCAM

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The ToO#7:

- First observation: October 27th 2014 as a single swath.
- Second observation: July 9th 2016 as a raster scan.

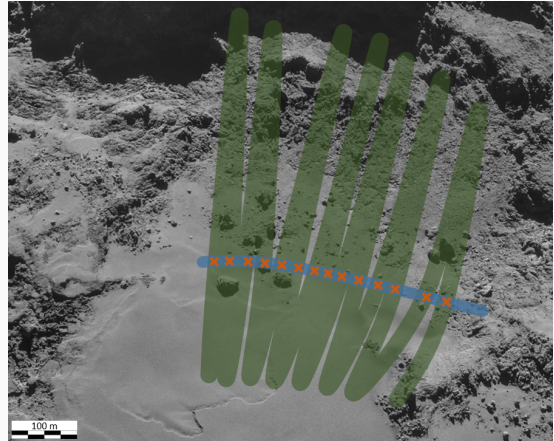


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The ToO#7:

- First observation: October 27th 2014 as a single swath.
- Second observation: July 9th 2016 as a raster scan.
- Objective: observe the same area twice with similar high spatial resolution.

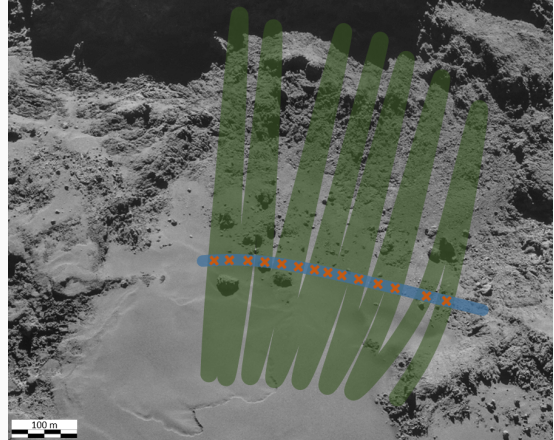


Figure 1: NAVCAM image of the Imhotep region with indicated the 2014 and 2016 swaths. Copyright: ESA/Rosetta/NAVCAM

Measurements

1. Context
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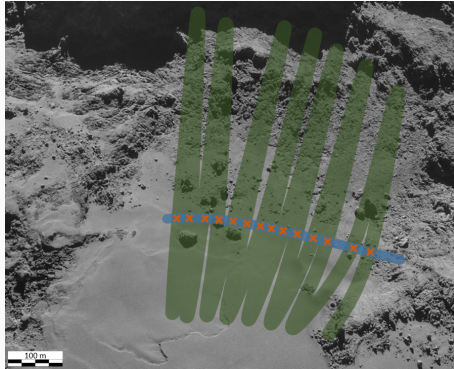


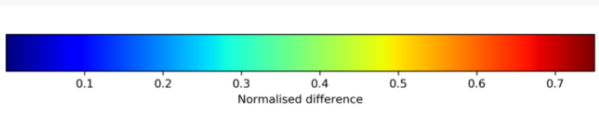
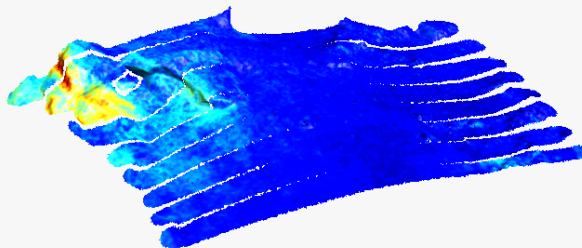
Figure 2: NAVCAM image of the Imhotep region with indicated the 2014 and 2016 swaths. Copyright: ESA/Rosetta/NAVCAM

Resolution	SMM	MM
October 2014	$\approx 20m$	$\approx 60m$
July 2016	$\approx 30m$	$\approx 90m$

The 2016 raster scan intersected the 2014 swath a total of 14 times.

What measurements to use ?

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Model of the subsurface

Thermal and radiative model

1. Context
2. Model of the subsurface
3. Results

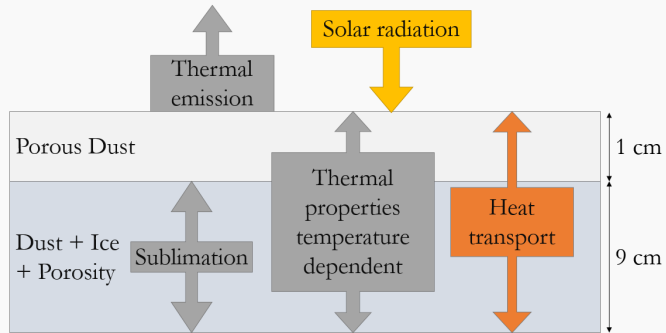


Figure 3: Simplified thermal model of the subsurface of 67P/C-G, indicating the processes at play

Thermal and radiative model

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Parameters	MM (188 GHz)	SMM (562 GHz)
Pure dust density (Kg/m^3)	2000	
Pure ice density (Kg/m^3)	925	
Emissivity	0.9	
Albedo	0.06	
Particulate size	400 μm	
Penetration depth	$\approx 4\text{cm}$	$\approx 1\text{cm}$
Upper layer porosity	80%	
Lower layer thermal conductivity	Variable	
Lower layer density		

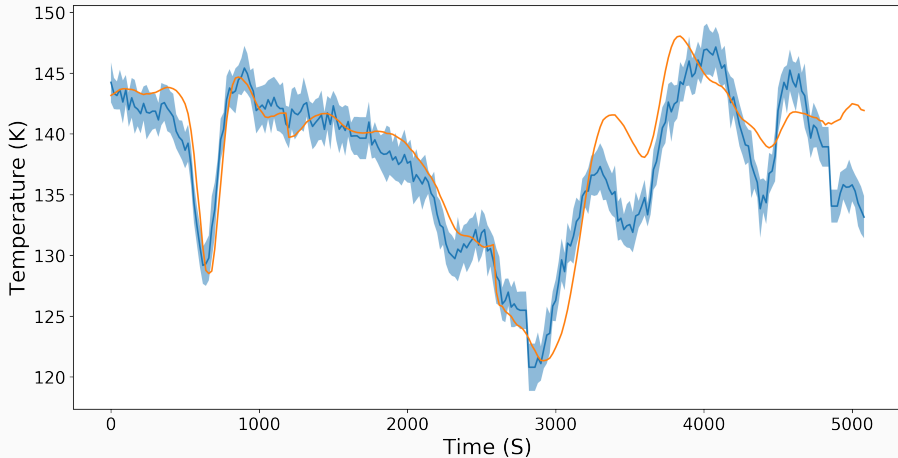




Results

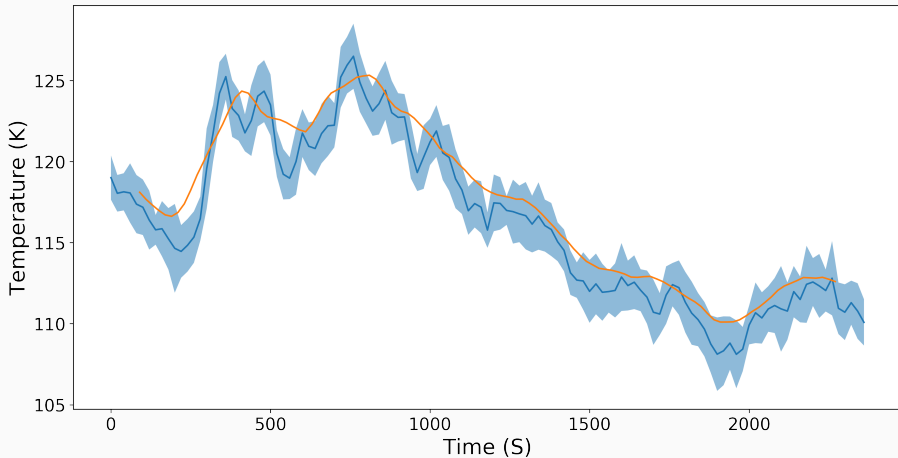
Fit of the model to the measurements - 2014

1. Context
2. Model of the subsurface
3. Results



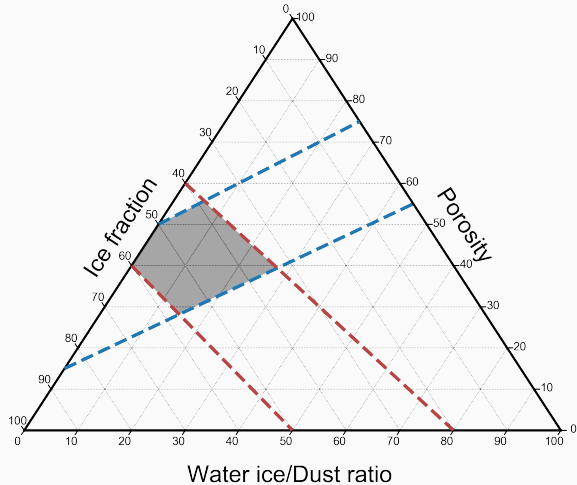
Fit of the model to the measurements - 2016

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2. Model of the subsurface
3. Results



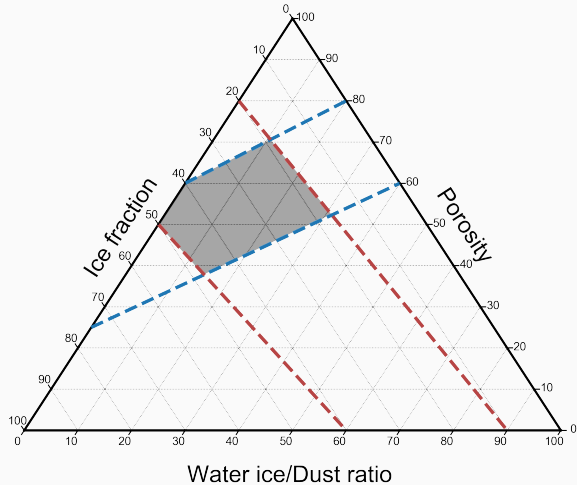
Constraints of the subsurface 2014

1. Context
2. Model of the subsurface
3. Results



Constraints of the subsurface 2016

1. Context
2. Model of the subsurface
3. Results



Changes in the subsurface properties

1. Context
2. Model of the subsurface
3. Results

Phase	2014	2016
Dust	0-28%	0-32%
Water ice	45-70%	30-60%
Vacuum	30-55%	40-70%



Conclusions

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- The fit of the model to the measurements can still be improved.



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- The fit of the model to the measurements can still be improved.
- Values presented and error bars are not final



Conclusions

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- The fit of the model to the measurements can still be improved.
- Values presented and error bars are not final
- Between the 2014 (before perihelion) and 2016 (after perihelion) the constraints on the subsurface shift towards lower water ice fraction and densities.



Perspectives

1. Context
2. Model of the subsurface
3. Results

- The model used in this work offers a powerful tool to investigate the subsurface of 67P/C-G using MIRO data and can help constrain the composition of the subsurface and its evolution through time.



Perspectives

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- The model used in this work offers a powerful tool to investigate the subsurface of 67P/C-G using MIRO data and can help constrain the composition of the subsurface and its evolution through time.
- More parameters need to be investigated in order to better fit the measurements of the Imhotep region measurements.



Perspectives

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3. Results

- The model used in this work offers a powerful tool to investigate the subsurface of 67P/C-G using MIRO data and can help constrain the composition of the subsurface and its evolution through time.
- More parameters need to be investigated in order to better fit the measurements of the Imhotep region measurements.
- MIRO made observations of Imhotep at other times, and these will be included in future studies.

